



gaia



Gaia News:Counting down to launch

A. Vallenari

INAF, Padova Astronomical Observatory
on behalf of DPAC

Outline

- Gaia Spacecraft status
- The Gaia sky
- Gaia open and globular clusters
- From data to science: data releases

Counting down to launch

- Gaia in Kourou on 23-26 August
- Launch campaign underway
- Payload Module is under electrical testing
- Sun Shield Deployment test 10-11 October
- S/C fuelling 2-5 Nov
- Launch 20 Nov 2013
08:57:30 UTC
- OR#4 successful
- Science performances





gaia

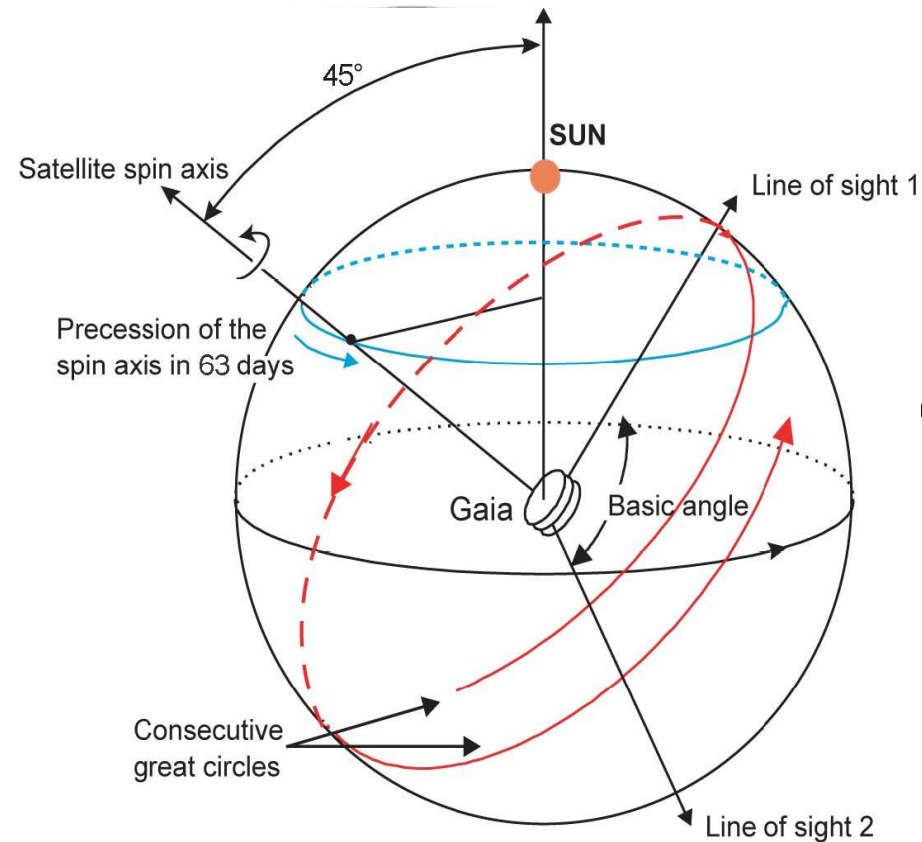
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Scanning the Galaxy

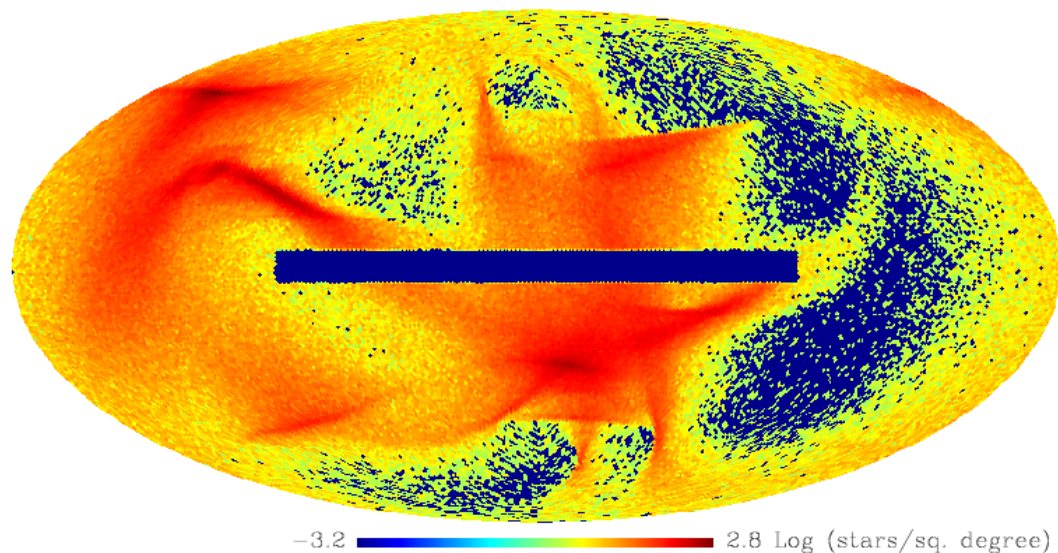
- Scanning Mission
- Complete down to $G=20$.
- Bright limit $G=5.7$
- Five-year scanning survey mission: [sky-coverage](#) non-uniformity
- Each source is observed ~ 75 times in astrometry & photom. 50 in spectroscopy
- Varies over the sky between
- ~ 50 and ~ 130 ($\sim 20\%$ dead time)
- Angular resolution comparable to HST



Astronomer's shopping list

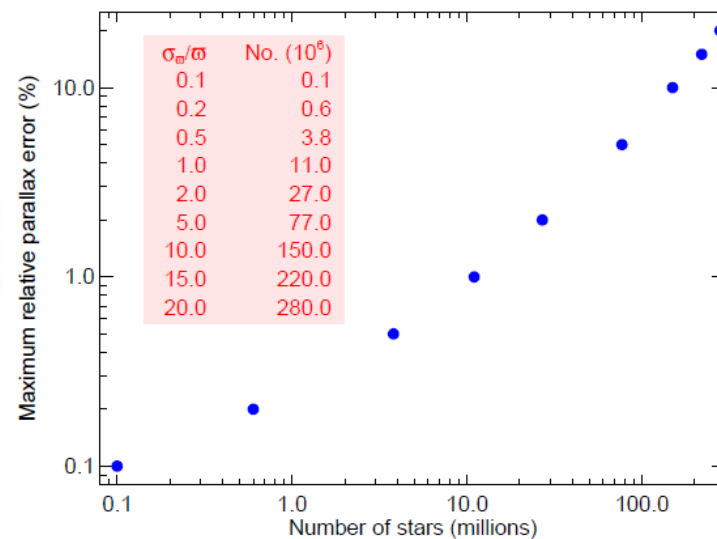
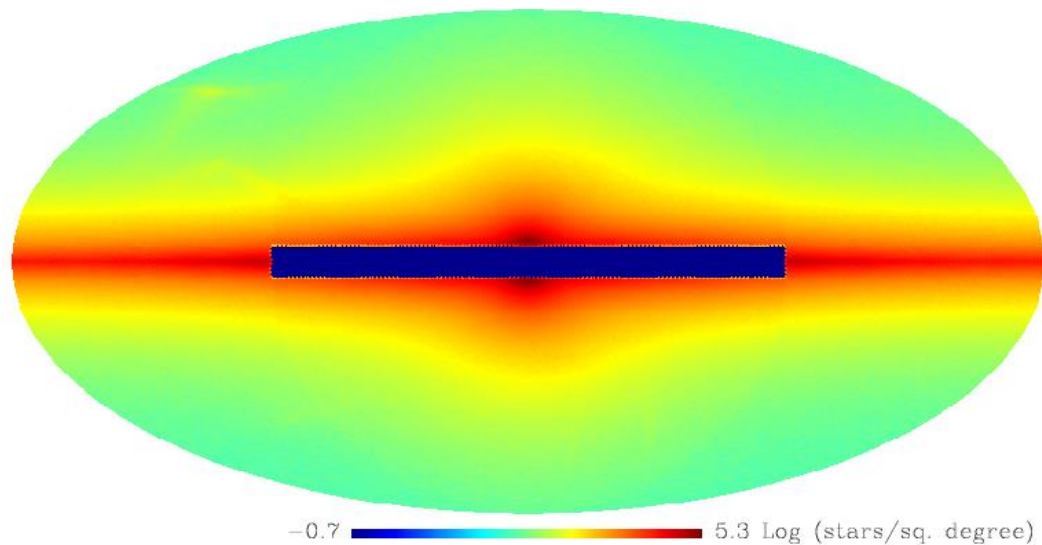
- High precision astrometry —————→ • **1 billion stars**
 - 10 μas @ $V < 13$ mag
 - 25 μas @ $V = 15$ mag
- High-precision broad-band photometry
- Multi-band photometry
- Spectra IR Ca II triplet
- Spectral classification ($\log g$, T_{eff} , $[\text{Fe}/\text{H}]$) (Bailer-Jones+2013)
 - Radial velocities —————→ **300millions of stars**
 - Metallicities —————→ **a few millions**
 - Rotation velocities
- Single Stars, but not only: SSO (Mignard+2012), binaries, galaxies(Bellas-Velidis2013), QSOs,exoplanets (Sozzetti+ 2008)

Gaia sky

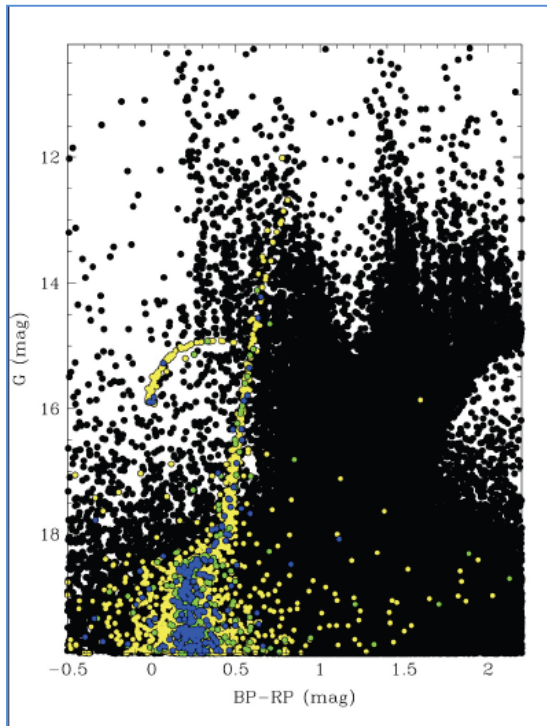


9E8 Thin Disk,
 4.3E8 Thick Disk stars,
 2.1E7 Halo stars
 1.7E8 Bulge stars

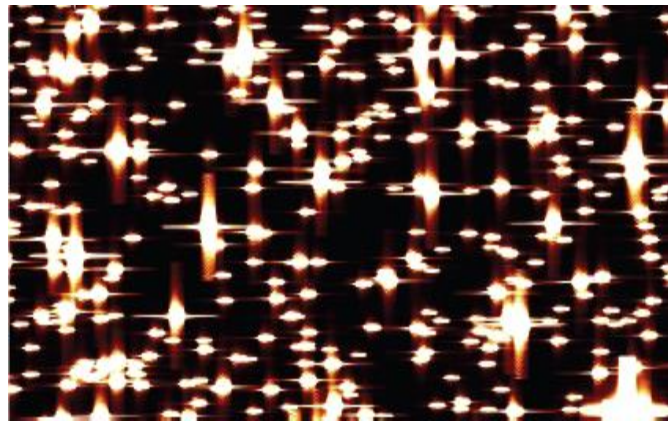
Colors: stellar density
 (X. Luri 2010)



$d=10$ Kpc, $c=2.5$, bulge

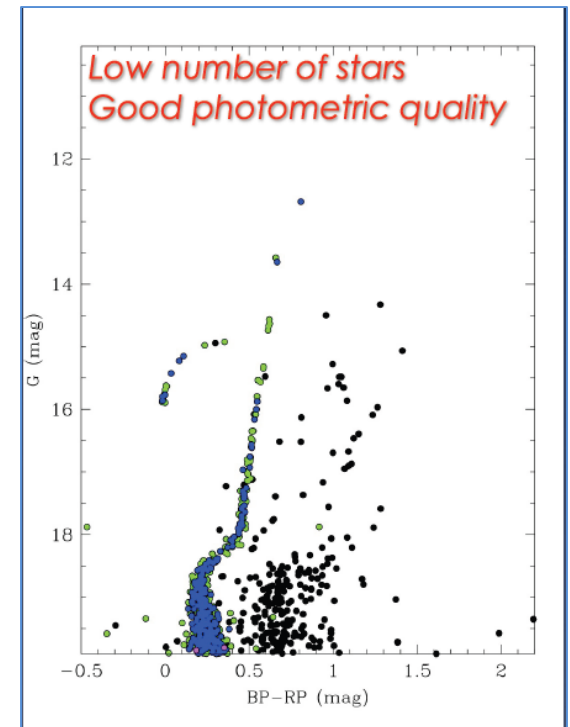


Gaia GIBIS FoV

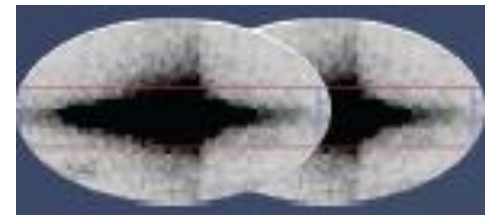


Pancino+2013

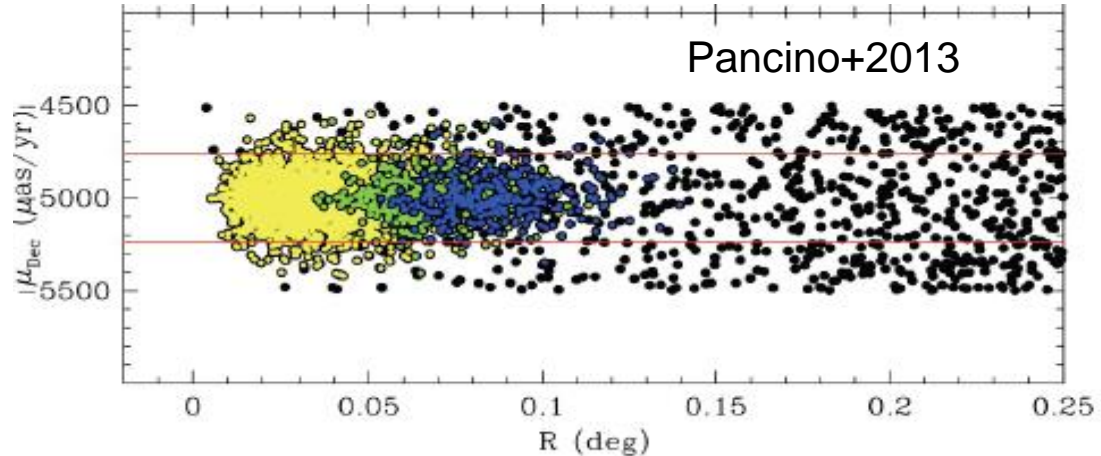
p.m selection



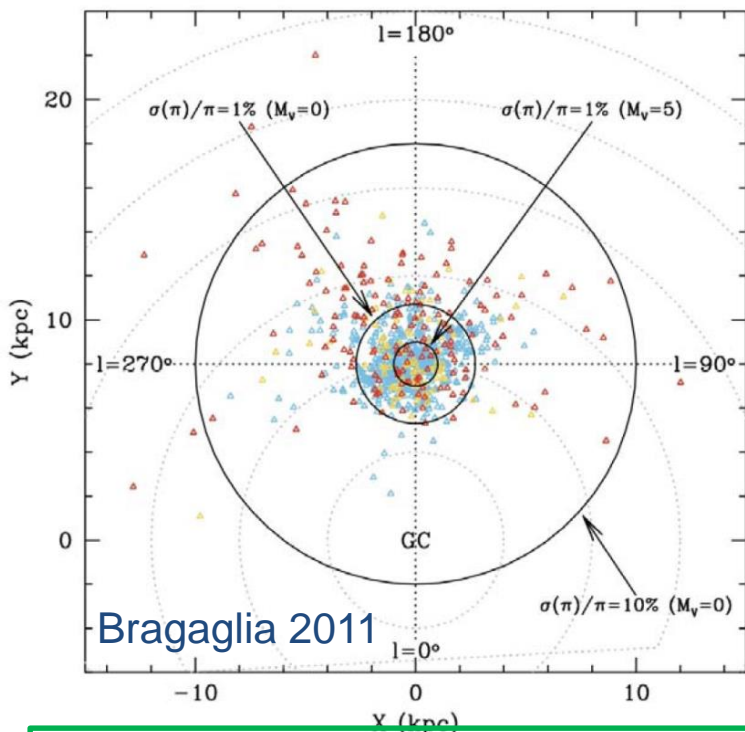
Non rotating 3D cluster (Kupper + 2011) +field stars
Generating Gaia images
crowding (partial superposition)



- 80 of 150 GCs
d < 10 Kpc
- Proper motions – orbits – maybe tidal tails → halo potential
- Membership determination of 100/10000 stars(outside the half light radius, 3-5 μ arcsec/yr at 10 Kpc)
- More than 5000 stars for half of the clusters
- Mean distances to < 1 % for about 80 clusters
- Mean distances to < 5 % for all clusters
- Spectra of stars above G=17 for Rv



Gaia Open Clusters-I



Present situation: 2095 known OCs
 1193 with distance
 100 with a [Fe/H] estimate (Dias+2010)

Gaia: Derive distances + pm of individual stars in Ocs

- at 1% for $M_v=5$ $d < 1.5 \text{ kpc}$
- at 1% for $M=0$ $d < 4 \text{ kpc}$
- at 10% for almost all known cluster

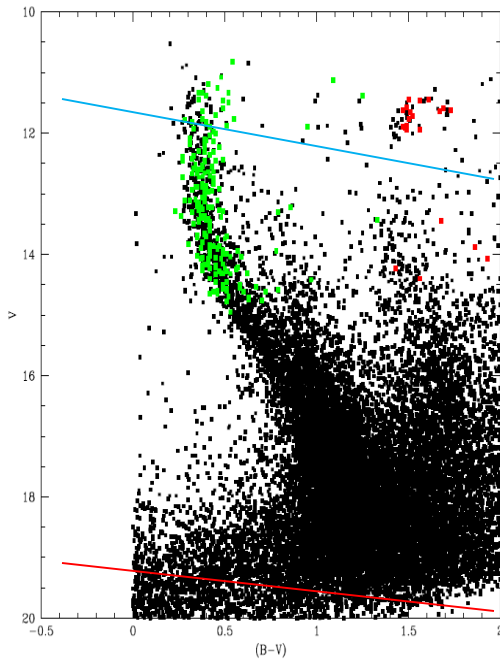
→ accurate membership-- orbits

Small velocity dispersion in OCs (1 - 2 km/sec) → studies of the internal dynamics require ~ 0.2 km/sec

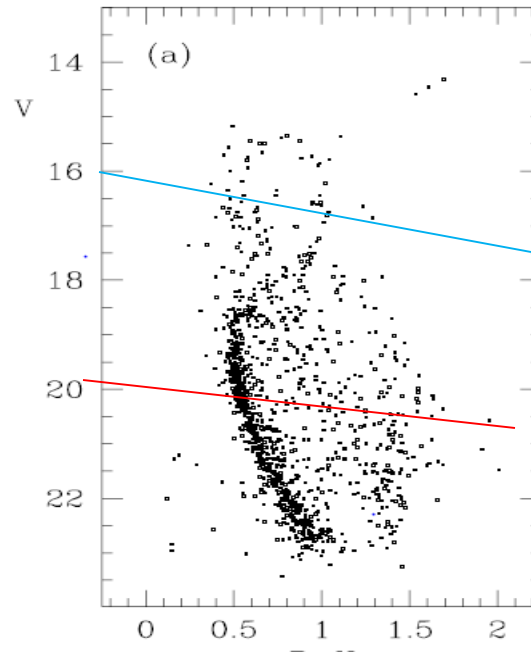
- Gaia: accuracy better than 1% for transverse velocity
- G0 stars brighter than $V \sim 13$ ($d < 500 \text{ pc}$), K1 III (red clump in old OCs) $V < 14$: $d < 5 \text{ kpc}$.

Missing: Detailed chemical abundances for $G > 11-12$

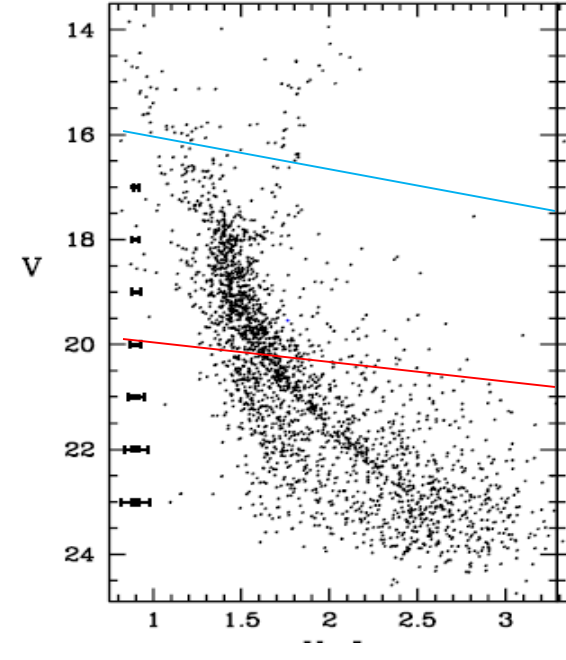
Gaia Open Clusters-II



B-V



V-I



B-V

NGC 6705

(Vallenari+2013)

D=1800 pc,

Age=250-300 Myr

Be 17

(Vallenari+1999,

Bragaglia+ 2006)

D=2600 pc,

Age=10 Gyr

Be 29 (Tosi,

Bragaglia 2006)

D=13.05 Kpc,

Age=3.7 Gyr

From data to science: Gaia Catalog

- Complex catalog dealing with more than 2 billions of objects, having information about: positions, proper motions, parallaxes, radial velocities, Aps, binarity, variability

- What will you get:

- Validated Gaia data
- Documentation
- Advanced tool for data manipulation, data mining, and visualization
- cross.-match with existing Catalogs and survey data



Catalog Validation

- Internal consistency checks
 - proper motion vs. distance
 - photometry vs. Spectroscopy
- Comparing to models and external data
 - detecting artifacts from the data processing
 - systematics in Gaia data
- Apply data mining techniques
 - identify outliers and unexpected correlations
 - identify and document expected correlations
- Special objects: clusters
- Cross method variability analysis
 - variability correlated with binary component separation
 - periodicities correlated with scanning law

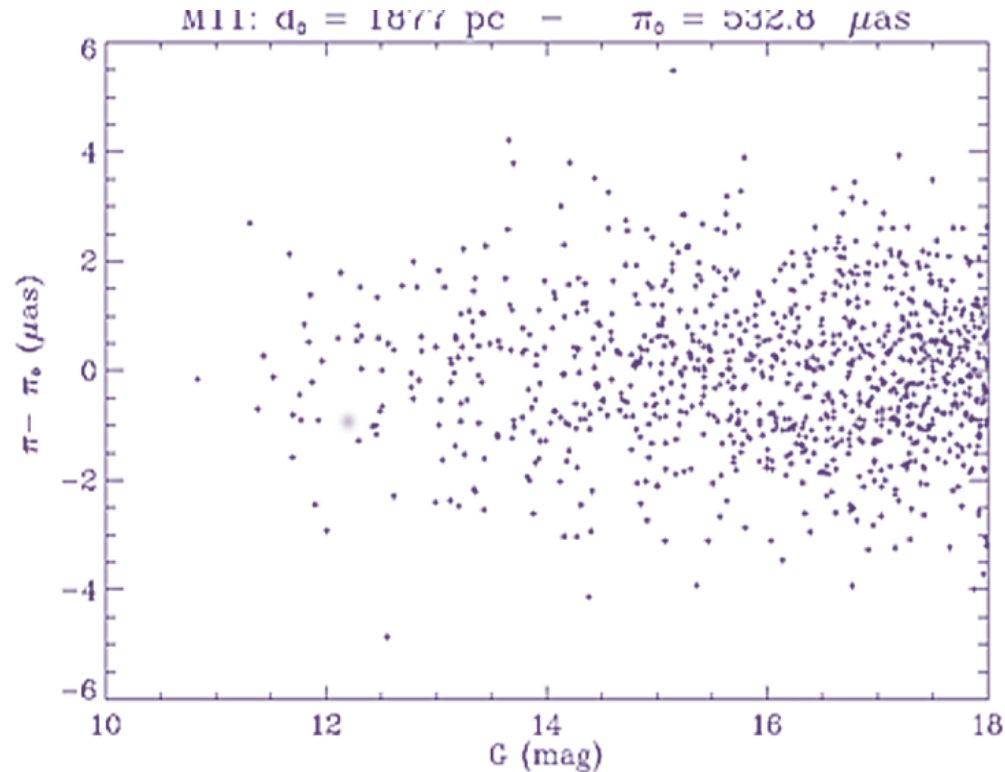
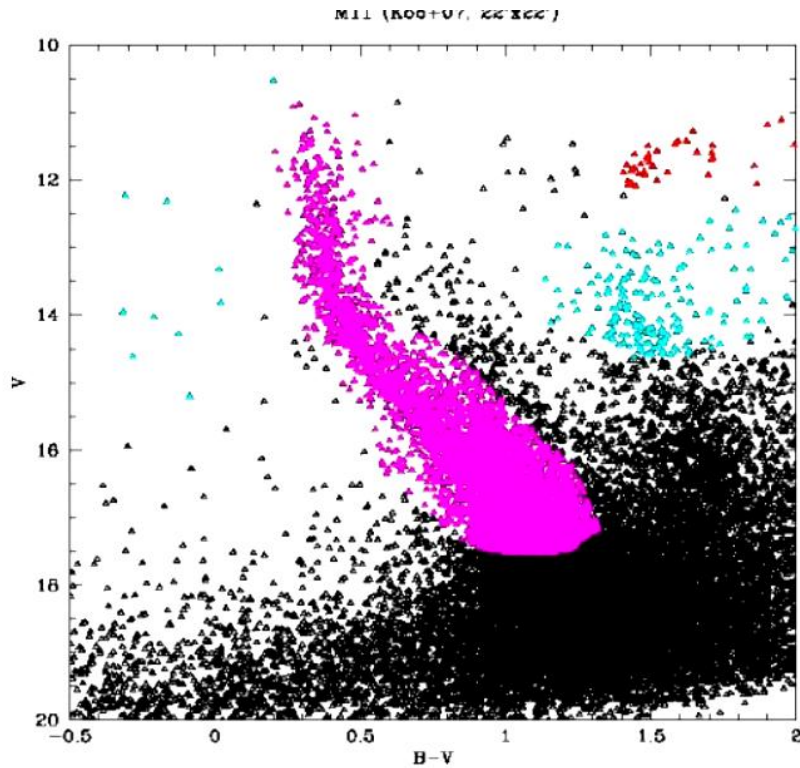


Using Ocs for Catalog Validation

- Ocs are single stellar population: same age and metallicity
- large samples (10^2 - 10^3) of stars having a 3D spatial spread of about **10 pc** and a 3D velocity dispersion **<1 km/s**.
- Ideal tests of Gaia astrometry, stellar parameters
- Cross match with external Catalogs+comparison with stellar models → ground based surveys
- Pleiades MS problem and 10% distance discrepancy: systematics over 1 deg in Hipparcos data? (van Leeuwen 2009)

Table 1. Pleiades parallaxes (updated from Soderblom *et al.* 2005)

Method	π_{abs} (mas)	D (pc)	m-M	Ref.
<i>Hipparcos</i> all-sky	8.45 ± 0.25	118.3 ± 3.5	5.37 ± 0.06	2
<i>Hipparcos</i> new reduction	8.18 ± 0.13	122.2 ± 1.9	5.44 ± 0.03	7
Main-sequence fitting	7.58 ± 0.14	131.9 ± 2.4	5.60 ± 0.04	1
Allegheny Observatory parallaxes	7.64 ± 0.43	130.9 ± 7.4	5.59 ± 0.11	3
Interferometric orbit	7.41 ± 0.11	135.0 ± 2.0	5.65 ± 0.03	4
Dynamical parallax	7.58 ± 0.11	131.9 ± 3.0	5.60 ± 0.05	5
<i>HST</i> FGS parallax of 3 Pleiads	7.43 ± 0.17	134.6 ± 3.1	5.65 ± 0.05	6



NGC 6705 (Vallenari+2013)
 D=1800 pc,
 Age=250 -300Myr
 Radius=6.2 pc
 Conservative Intrinsic d
 dispersion= 5 pc

Intrinsic $\sigma\pi=1.4\mu$ as
 << expected Gaia uncertainty at $G>11$
 To be used up to 2-3 Kpc
 (Spagna 2013)

Gaia Data: Why do you have to wait so long

- Gaia is self-calibrating
 - iterative process to derive astrometric parameters, attitude and stellar parameters
- Gaia principles involve global astrometry: no immediate scientific data from single observations
 - Total commissioning phase: 4-6 months
 - at least one full sky coverage needed for an astrometric solution (positions) → at [least 6 months of data](#)
 - at least 18 months of data for a full astrometric solution but sampling might be not sufficient in many cases
- Colors must be known to achieve good accuracy → calibrated photometry
- Each data release requires time to go from DPAC internal database to public archive (3 Months)

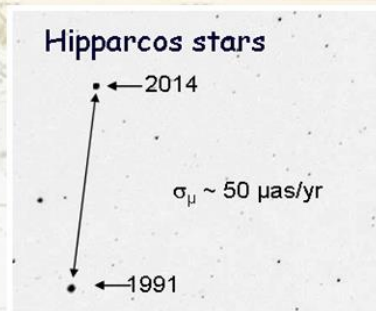
Data release scenario

Science alert: ASAP

Release

L + 22m

α deg 8	δ deg 9
4.18700521	41.06905721
57.84982093	-25.93298679
87.24506095	63.69702032

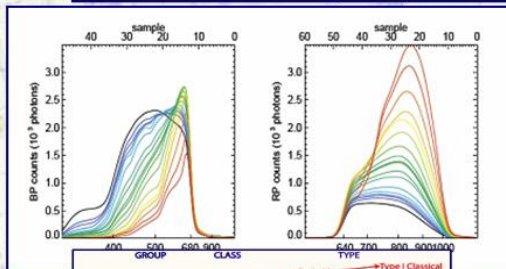


L + 28m

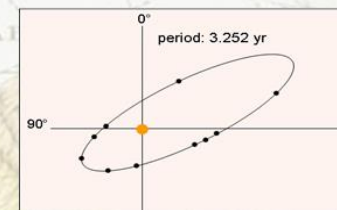
α deg 8	δ deg 9	π mas 11	μ_α^* mas/yr 12	μ_δ mas/yr 13
4.18700521	41.06905721	18.92	190.14	27.73
57.84982093	-25.93298679	21.42	273.72	167.93
87.24506095	63.69702032	4.12	5.07	-98.19
97.98432970	-27.72950422	17.95	21.31	217.87
106.43737563	26.14031575	21.74	18.58	-10.54
109.41026120	-29.01434479	7.92	-4.53	-7.01

G **G_{BP}** **G_{RP}** **V_r**

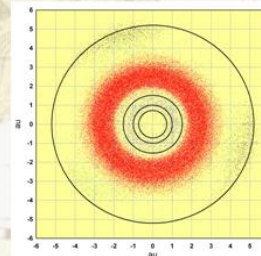
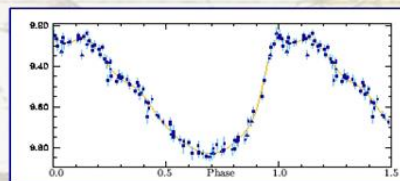
L + 40m



$\langle V_r \rangle$



L + 65m



Clusters in H T P M catalog

+ accuracy on pm (**best case**) : 30 to 250 $\mu\text{as/yr}$, depending on the magnitude, with an average of 50 $\mu\text{as/yr}$ (20 times better than the average 1000 $\mu\text{as/yr}$ of the Hipparcos values)

- + All OCs closer than 300 pc + the richest OCs up to 500 pc (19 objects, vanLeeuwen 2009)
- + 150 stars in Hyades to 40-10 stars in distant OCs ($V=12$)
- + 80 more candidates from Francis +2012
- + Full sample of clusters in the second release+

- Gaia is approaching launch
- Final tests are going on
- DPAC is ready to deliver the promises of Gaia
- A paper will be published after commissioning to update Perryman 2001

