Wavefront Reconstruction for Imperfect Pyramid Wavefront Sensor Assemblies: Generalizing the controller slope space

Or: A workaround to WFS alignment constraints.

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Observatoire de Paris - LESIA

Wavefront Sensing in the VLT/ELT era II, Oct 2^{nd} - 4^{th} 2017







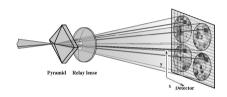
- Pyramid Wavefront Sensor recap.
- 2 Mishaps: prism defects and misalignments
- Secondary Space Control
 Secondary Space Control
- Misalignment transfer function: Fourier domain loss analysis
- 5 Bench and end-to-end simulations: ESC performance
- 6 Conclusion

Introduction

 Focal plane WFS concept by R. Ragazzoni (1996)

Sensitivity increase over SH WFSs:

- Telescope instead of microlens diffraction limit. >80x gain for ELTs!
- Modulation: user-selectable sensitivity knob
- Hot topic in instrumental developments: Keck, Subaru, TMT, GMT, E-ELT (HARMONI, METIS, MICADO)



PWFS technology is maturating:

- Calibration, drifts, nonlinearities
- NCPAs, Optical gain, ...
- Theoretical understanding
- Misalignments and defects

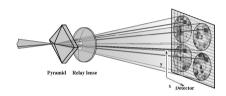


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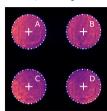
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What are Pyramid signals?

PWFS - Quadrant registration



$$n = \langle A + B + C + D \rangle_{(x,y)}$$

$$\longrightarrow S_x = \frac{A - B + C - D}{n}(x,y) \longrightarrow$$

$$S_y = \frac{A + B - C - D}{n}(x,y)$$

 S_x, S_y slopes map for the reference point.

Y-axis slopes

X-axis slopes

How do we interpret these slopes maps?

Ragazzoni 96: Ray optics – Modulation-tuned gradient sensor with neat saturation.

Vérinaud 04: 1-D derivations – gradient/phase linear sensor dep. on frequency range.

Fauvarque 16: The PWFS actually has an OTF, and so do the slopes map S_x and S_y . Convolutional reconstructors are possible.

What are Pyramid signals?

What is known on slope signals S_x and S_y :

- Direction-sensitive operators.
- Linear operators in small-phase/closed-loop regime.
- Somewhere between phase and gradient within frequency range.
- Theoretically permit phase reconstruction.

When the PWFS is misaligned: distorted S_x , S_y are measured. Distorted S_x , S_y may not contain all phase information.

Key goal: ensure measurements always contain enough information:

- to achieve valid wavefront reconstruction
- to operate a stable, well-conditioned AO loop.



Misfabrications and misalignments

Many possible prism fabrication errors cause:

- Zero point quadrant flux variations
- Non-square quadrant layout

Theoretical perfect PWFS requires:

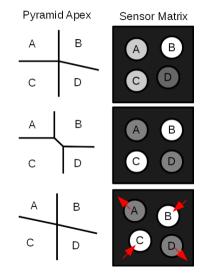
- Perfect rectangle layout
- Identical quadrant flux
- Integer pixel spacing between quadrants

A, B, C, D pixels must match exactly for PWFS validity.

How tight is specification: 1/10th pixel?

Software quadrant fit and center select: $1/2~{\rm px.}$ guaranteed. Hardware accelerated processing: offset may be larger.

 \rightarrow Impact study of free translations of all 4 quadrants.



0.5 Pixel Misalignments

Simulated 18 m telescope with 40x40 SCAO, r₀ 12.9 cm (more in Results section)

Perfect quadrant alignment

Two quadrants $0.5~\mathrm{px}$ offset

Invisible speckle zones - Wavefront residual 150 ightarrow 196 nm RMS: extra 125 nm RMS !

Consequences for AO loops

Distorted PSFs even for low misalignments (< 0.5 px.) Possible solutions:

- Let distorted PSFs happen, ignore them or deconvolve data
- Mitigate the effect by filtering out some of high order modes (but actuators are expensive).
- Specify system to have close-to-perfect alignment



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- Specify system to have close-to-perfect alignment
- Do not only use $[S_x, S_y]$ control!



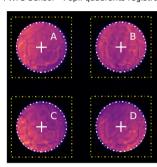
Expanded Space control: double the slope space.

Traditional gradient control slopes:

$$\begin{bmatrix} S_x \\ S_y \end{bmatrix} (x, y) = \begin{bmatrix} 1 & -1 & 1 & -1 \\ 1 & 1 & -1 & -1 \end{bmatrix} \cdot \begin{bmatrix} A \\ B \\ C \\ D \end{bmatrix} (x, y)$$

Expanded slope space:

PWFS Sensor - Pupil quadrants registration



PYRCADO bench: registration of quadrants on PWFS camera sensor for classical gradient control (white) and for ESC (yellow)

 $S_x^{\rm True}$ and $S_y^{\rm True}$ allow full reconstructor. In the misaligned case:

$$\begin{bmatrix} A(x,y) \\ B(x,y) \\ C(x,y) \\ D(x,y) \end{bmatrix} = \begin{bmatrix} A^{\operatorname{True}}(x-x_A,y-y_A) \\ B^{\operatorname{True}}(x-x_B,y-y_B) \\ C^{\operatorname{True}}(x-x_C,y-y_C) \\ D^{\operatorname{True}}(x-x_D,y-y_D) \end{bmatrix}$$

 S_{ν}^{True} and S_{ν}^{True} allow full reconstructor. In the misaligned case:

PWFS mishaps

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In the Fourier space:

$$\begin{bmatrix} A(x,y) \\ B(x,y) \\ C(x,y) \\ D(x,y) \end{bmatrix} = \begin{bmatrix} A^{\mathrm{True}}(x-x_A,y-y_A) \\ B^{\mathrm{True}}(x-x_B,y-y_B) \\ C^{\mathrm{True}}(x-x_C,y-y_C) \\ D^{\mathrm{True}}(x-x_C,y-y_D) \end{bmatrix} \qquad \begin{bmatrix} \tilde{A}(u,v) \\ \tilde{B}(u,v) \\ \tilde{C}(u,v) \\ \tilde{D}(u,v) \end{bmatrix} = \begin{bmatrix} e^{2i\pi(x_Au+y_Av)}\tilde{A}^{\mathrm{True}}(u,v) \\ e^{2i\pi(x_Bu+y_Bv)}\tilde{B}^{\mathrm{True}}(u,v) \\ e^{2i\pi(x_Cu+y_Cv)}\tilde{C}^{\mathrm{True}}(u,v) \\ e^{2i\pi(x_Du+y_Dv)}\tilde{D}^{\mathrm{True}}(u,v) \end{bmatrix}$$

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with the **P** transform between [A, B, C, D] and $[S_x, S_y, S_z, S_f]$:

$$\begin{bmatrix} \tilde{S}_{\mathbf{T}^{\mathrm{True}}}^{\mathrm{True}} \\ \tilde{S}_{\mathbf{T}^{\mathrm{True}}}^{\mathrm{True}} \\ \tilde{S}_{\mathbf{f}^{\mathrm{True}}}^{\mathrm{True}} \end{bmatrix} (u, v) = \underbrace{\frac{1}{4} \mathbf{P} \overline{\Delta(u, v)} \mathbf{P}^{\mathsf{T}}}_{\mathsf{Mis}} \begin{bmatrix} \tilde{S}_{x} \\ \tilde{S}_{y} \\ \tilde{S}_{z} \\ \tilde{S}_{f} \end{bmatrix} (u, v) \text{ with } \Delta(u, v) = \mathrm{Diag} \begin{bmatrix} e^{2i\pi(x_{A}u + y_{A}v)} \\ e^{2i\pi(x_{B}u + y_{B}v)} \\ e^{2i\pi(x_{C}u + y_{C}v)} \\ e^{2i\pi(x_{D}u + y_{D}v)} \end{bmatrix}$$

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We now have a transfer function between $[S_x, S_y, S_z, S_f]$ and $[S_x^{\text{True}}, S_y^{\text{True}}]$



The Mis transform: An example

Misalignment test case:

$$x_A, y_A = -0.24, +0.46$$

 $x_B, y_B = +0.28, -0.49$
 $x_C, y_C = -0.17, +0.38$
 $x_D, y_D = +0.45, -0.47$

• All offsets ≤ 0.5 pixels

For pupils of 55 px with 100 px separation, is equivalent to:

- 2% tol. in deviation angle
- 12 mrad rotation.



The **Mis** transform: An example

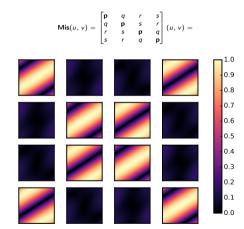
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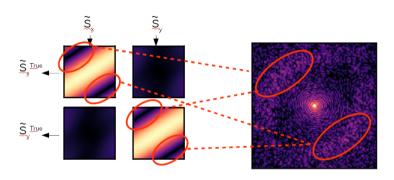
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Frequency dependence of all 16 terms of the Mis transform

The **Mis** transform with $[S_x, S_y]$ control



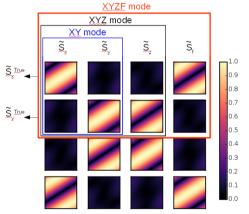
Speckles remain due to unseen frequencies!

- Frequencies coupled in $[S_z, S_f]$ are missing from sensor slopes $[S_x, S_y]$.
- $[S_x^{\text{True}}, S_v^{\text{True}}]$ cannot be reconstructed.
- Reconstructor has bad SNR / is rank-deficient.



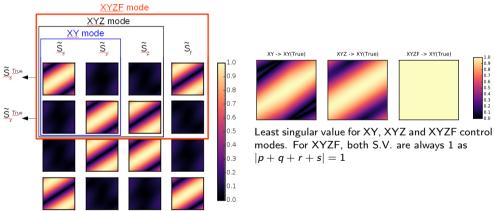
Three Control Modes

Can use $[S_x, S_y]$, $[S_x, S_y, S_z]$ or $[S_x, S_y, S_z, S_f]$. What is the loss ? We need a figure of merit for control modes. $[S_{\bullet}...] \longrightarrow [S_x^{\mathrm{True}}, S_y^{\mathrm{True}}]$



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Expanded Space Control Performance:
Bench and Simulation Results

Comparing Control Modes - Simulation Results

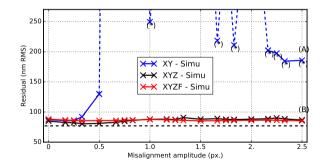
End-to-end simulation wavefront residuals (fitting removed) for 3 control modes. Misalignments from 0 to 2.5 subapertures.

- 18 m Telescope
- 40×40 DM
- 52×52 PWFS @ 658 nm, $r_{\text{Mod}} = 8\lambda/D$
- Kolmogorov ground layer, $r_0 = 13$ cm, v = 10 m.s⁻¹

Analytical error budget:

Fitting: 125 nm RMS (subtracted)

Rest: 77 nm RMS



Comparing Control Modes - Simulation Results

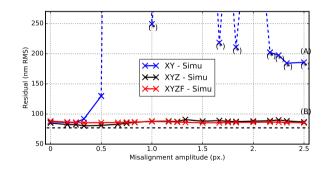
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XY explodes for < 0.5 px shifts. XYZ and XYZF modes remain nominal.

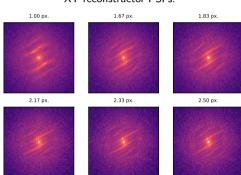
(*) data: Loop stability at high offsets

For high offsets, loop may still be stable: (\star) data points at [180, 250] nm RMS. Artifacts on long exposure PSFs illustrate **Mis** reconstructor blind zones:

Nominal H-band PSF (XYZF, 2.5 px offsets):



XY reconstructor PSFs:



Comparing Control Modes - Optical Bench

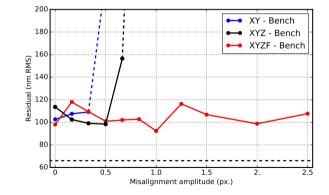
Optical bench: wavefront residuals (fitting removed) for 3 control modes. Misalignments from 0 to 2.5 subapertures.

- SLM with 500 px in pupil (virtual DM, turbulence screen)
- 18 m Telescope: 3.6 cm.px⁻¹
- 40×40 DM
- 260×260 PWFS @ 658 nm, $r_{\text{Mod}} : 8\lambda/D$. Rebin: 44×44 subap.
- Kolmogorov ground layer, $r_0 = 13$ cm, v = 10 m.s⁻¹

Analytical error budget:

Fitting: 125 nm RMS (subtracted)

Rest: 66 nm RMS



Comparing Control Modes - Optical Bench

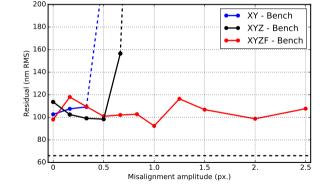
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Rest: 66 nm RMS



XY and XYZ are unstable for offsets > 0.5 px.

XYZF yields consistent performance regardless of alignment quality.

XYZ control: Why are both S_z and S_f required on the bench?

With perfect alignment and uniform, time-invariant pupil illumination:

$$S_z = \frac{A - B - C + D}{\langle A + B + C + D \rangle_{x,y}}(x,y) = 0$$
$$S_f = \frac{A + B + C + D}{\langle A + B + C + D \rangle_{x,y}}(x,y) = 1$$

$$S_f = \frac{1}{\langle A+B+C+D \rangle_{x,y}} (x,y) = 1$$

contain no information! (cf. Fauvarque 2016)

Simulation: misalignment, but illumination is OK. One of $[S_z, S_f]$ seems to be sufficient.

Bench: misalignment and changing flux

(pupil-plane speckles, non-conjugated aberrations, interference fringes, SLM artifacts, ...)

Both S_z and S_f are required!

Conclusion

On misalignments:

- Analyze effects of free quadrant translations on the sensor matrix.
- Demonstrate that gradient control is insufficient for imperfect situations due to misaligned pixel-to-pixel summations.
- Propose a figure of merit to quantify misalignment impact in gradient control.

Controlling the PWFS:

- We propose a control method for PWFS: Expanded Space Control
- ESC adds two 'slopes' maps in addition to gradients: cross-term S_z and the flux term S_f .
- **Simulation:** one of S_z , S_f makes the PWFS insensitive to misalignments.
- On bench: using both S_z , S_f required to maintain nominal performance.
- No loss is found from using extra S_z , S_f with PWFS well aligned.



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Gives: Very relaxed PWFS alignment constraints.

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Advertisement: COMPASS is now on GitHub, w/ all user-level code in Python 3.6!

Users & Contributors welcome!

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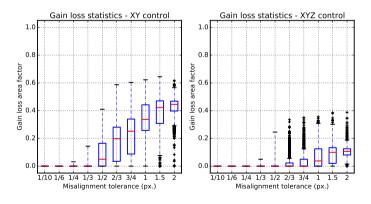
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Thanks for listening!

The loss metric: a Monte-Carlo analysis

Chosen metric: area of Nyquist domain for which least S.V. < 1/2.51 Fraction of modes with SNR loss > 1 magnitude



For 1/2 px. roundoff alignment - 50% chance of >5% loss at least with gradient control!

