

Osservatorio Astronomico di Brera

**Searching for absorbed-AGN
in the 2XMM-Newton Source Catalogue**

Elisabetta Memola

in collaboration with

R. Della Ceca, A. Caccianiga, P. Severgnini,
F. Cocchia, T. Maccacaro

**Part of an extended evaluation of
the 2XMM Catalogue in the context of the
Survey Science Centre work**

Scientific Goals

Obscured AGN should numerically dominate the AGN population

➔ 85% of the accretion power in the Universe could be absorbed (Fabian et al. 1998, MNRAS, 297, L11)

★ To select rare and interesting (e.g. Absorbed AGN, Type 2 QSOs) populations of X-ray sources to increase the current statistics.

★ Here we present a pilot-project with the aim of assembling a larger (about hundred) sample of obscured AGN

Absorbed AGN

To **efficiently** select absorbed AGN is not an easy task !

Nuclear properties affected by absorption in optical & soft X
➡ optical and soft X-ray surveys not efficient to select them

Hard X-ray data (2-10 keV) are needed
➡ Problem of absorption less severe

However at bright fluxes, even in the 4.5-7.5 keV band,
the majority of sources are 'unabsorbed' AGN
(75% - Della Ceca et al. 2006 from the Bright Survey)

Therefore efficient selection criteria are needed !

Selected results from the Hard Bright Sample

Selection Band = 4.5 - 7.5 keV, $F_x > 7 \cdot 10^{14}$ cgs

The nice correlation between the intrinsic absorption and HR2 found out for the HBS is very useful to exploit the incoming 2XMM catalogue

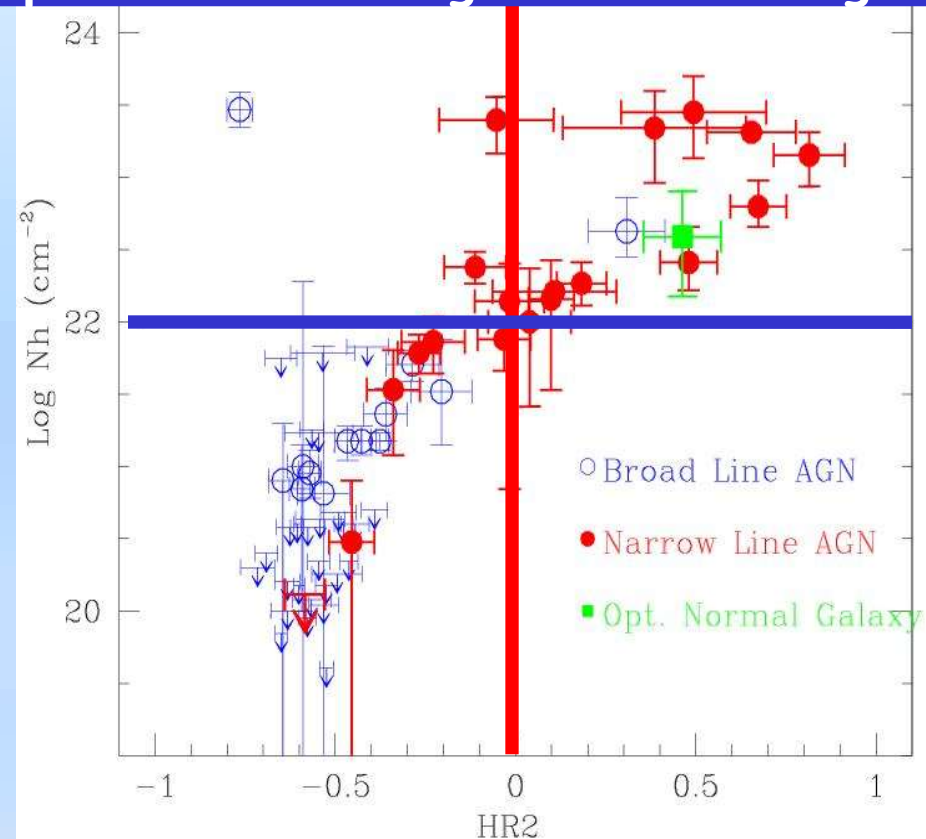
HR2 extremely efficient to select absorbed sources



100% of the sources with $HR2 > 0$ (up to $z \approx 0.8$) are **absorbed** and **4/13** are Type 2 QSOs

Energy bands:
2) 0.5 - 2.0 keV
3) 2.0 - 4.5 keV

$$HR2 = \frac{CR3 - CR2}{CR3 + CR2}$$



Caccianiga et al., 2004, *A&A*, 416, 901
Della Ceca et al., 2004, *A&A*, 428, 383
Della Ceca et al., 2005, astro-ph/0510845

Searching "with high efficiency" for absorbed AGN in the 2XMM catalogue (Watson et al., 2006, in preparation)

★ OABrera is a member of the XMM-Newton Survey Science Centre ★

Preliminary version (100.000 Sources)

Selections:

$|b| > 20$ deg

Exposure > 20 Ksec

Flux (4.5-12 keV) $> 10^{-13}$ cgs

Likelihood (4.5-12 keV) > 20

Target excluded

389 Sources

Work in progress @OABrera

Energy bands:

2) 0.5 - 2.0 keV

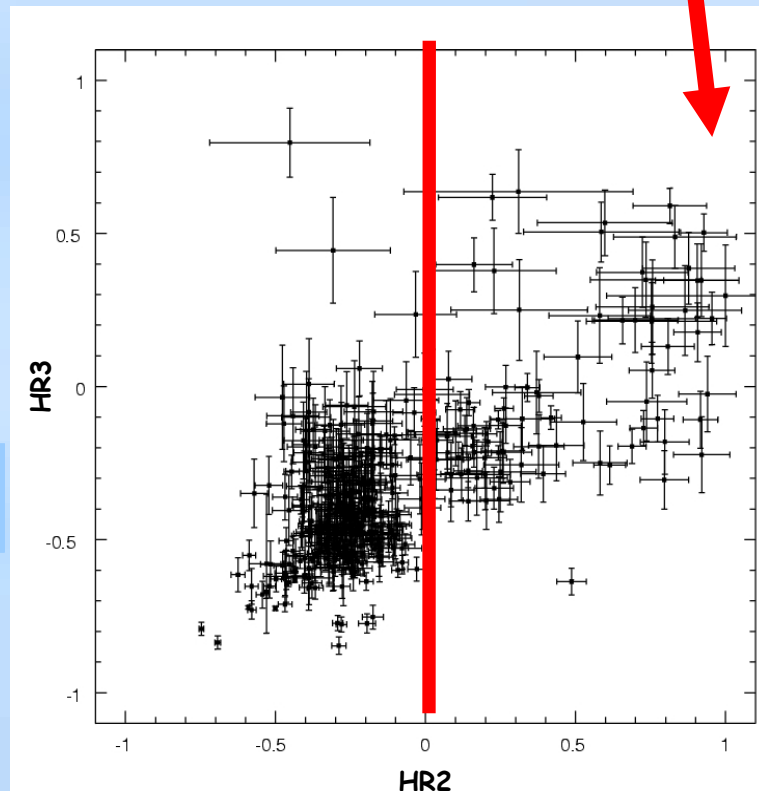
3) 2.0 - 4.5 keV

4) 4.5 - 7.5 keV

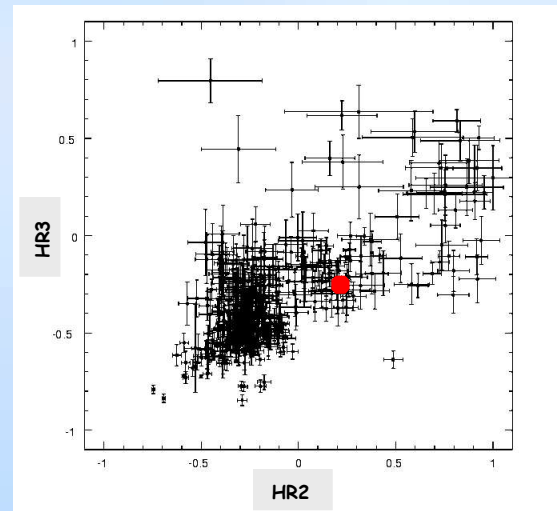
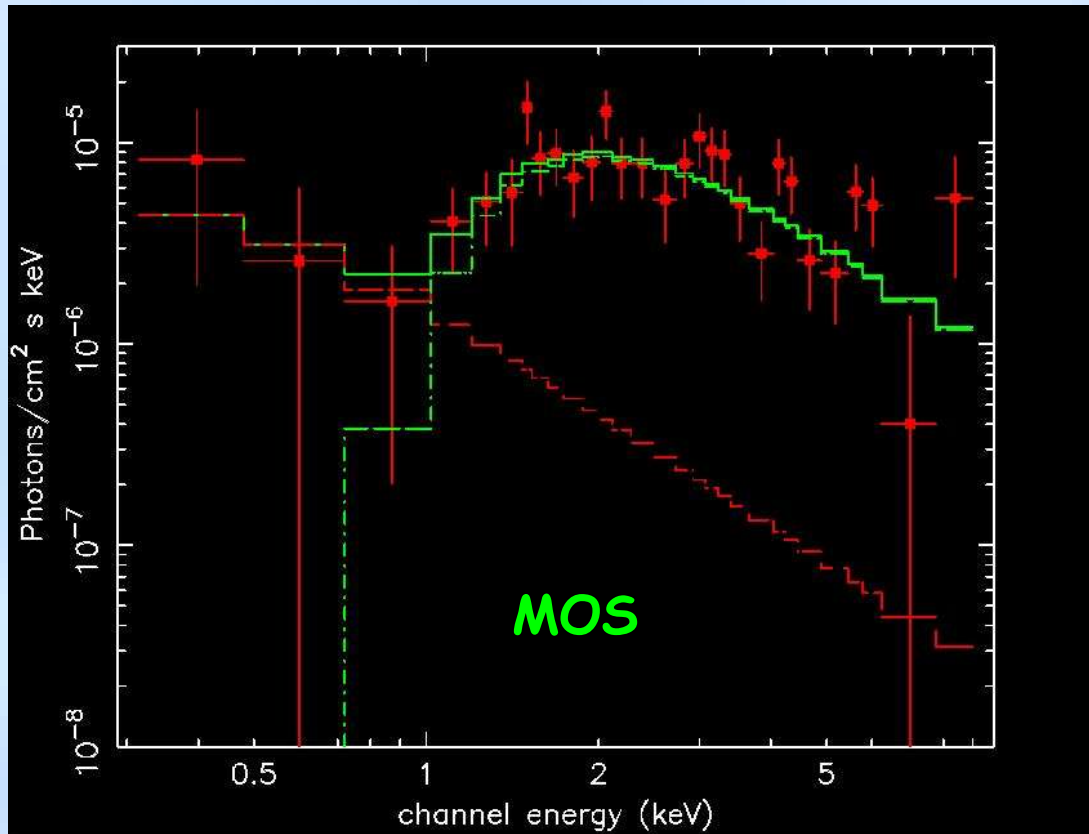
$$HR2 = \frac{CR3 - CR2}{CR3 + CR2}$$

$$HR3 = \frac{CR4 - CR3}{CR4 + CR3}$$

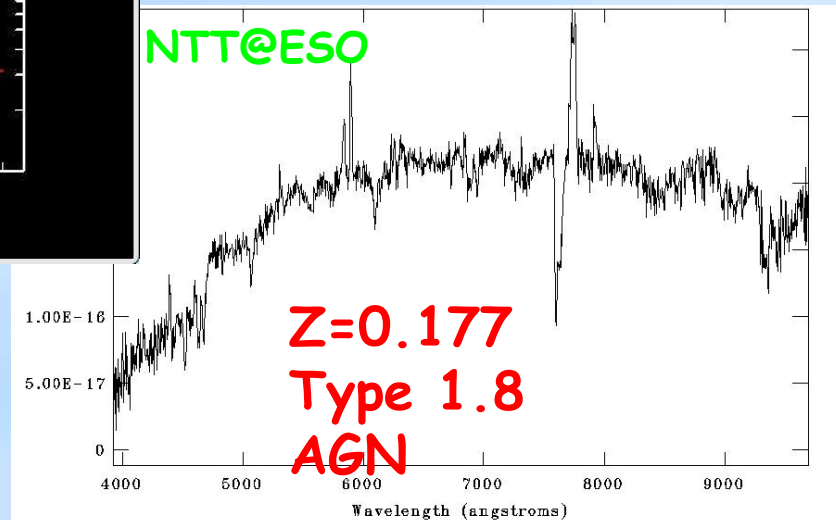
91 X-ray sources
with $HR2 > 0$



Newly identified absorbed AGN @OABrera



NTT@ESO



XMM X-ray spectrum

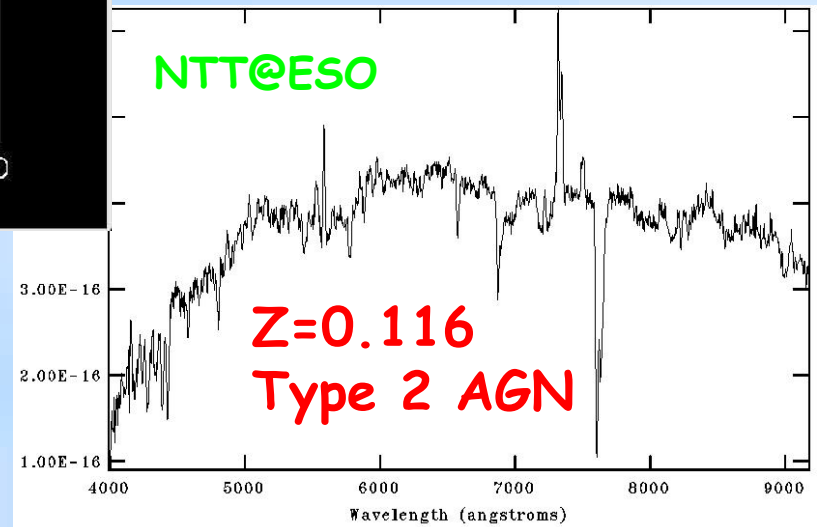
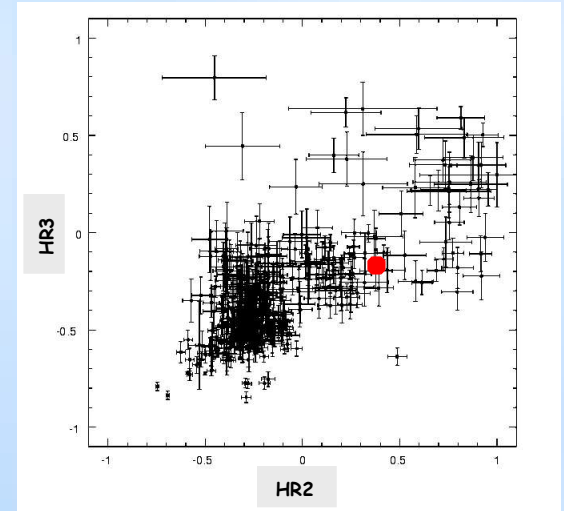
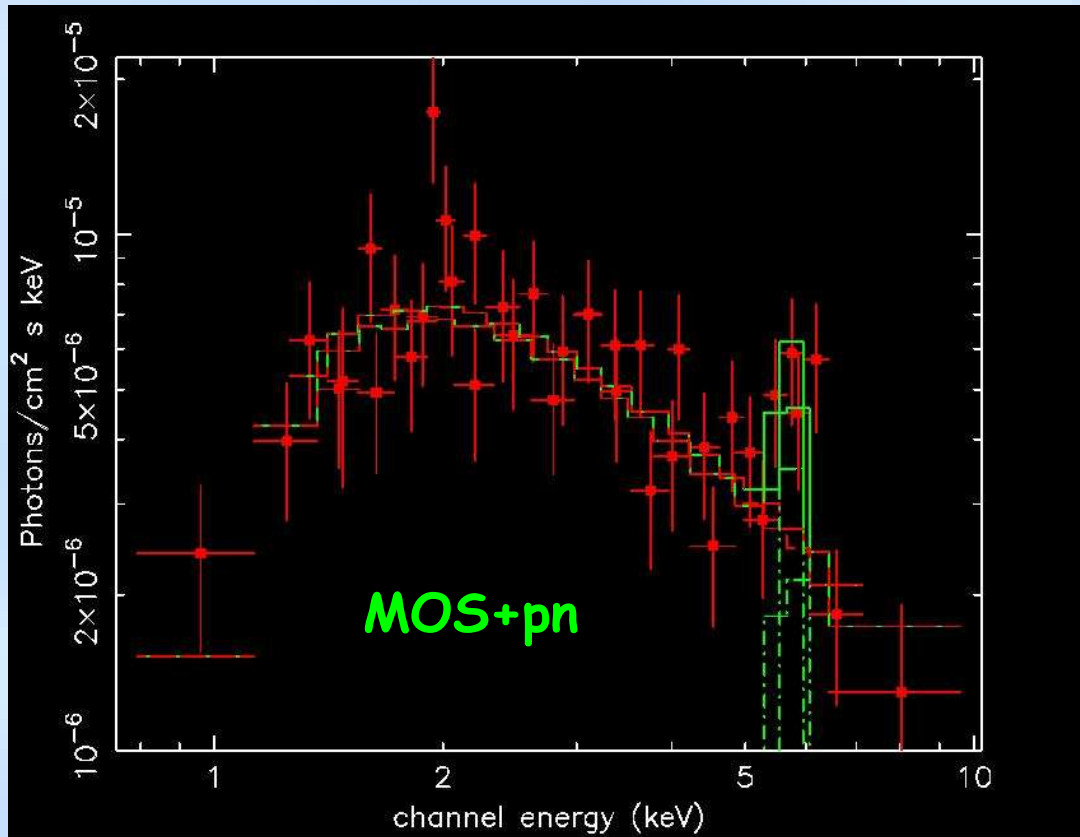
$$\Gamma \approx 1.9 \pm 0.5$$

$$N_H \approx 2.5 \pm 1.5 \times 10^{22} \text{ cm}^{-2}$$

$$L_{\text{Intr}(2-10 \text{ keV})} \approx 2 \times 10^{43} \text{ cgs}$$

Memola et al. 2006, in preparation

Newly identified absorbed AGN @OABrera



XMM X-ray spectrum

$$\Gamma \approx 1.4 \pm 0.3$$

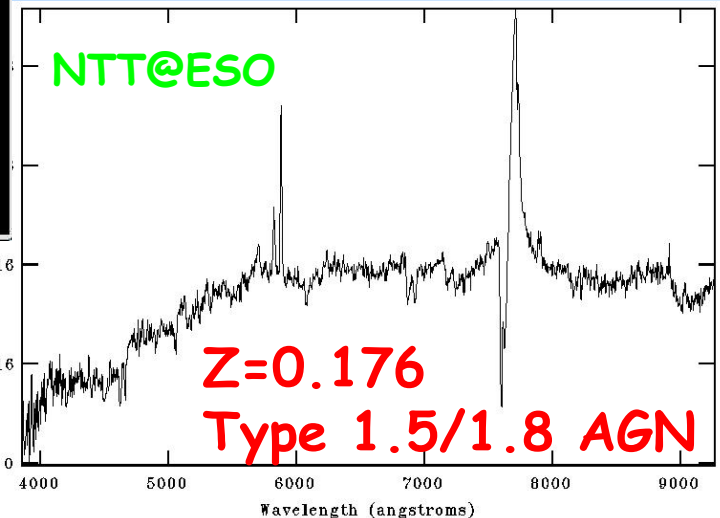
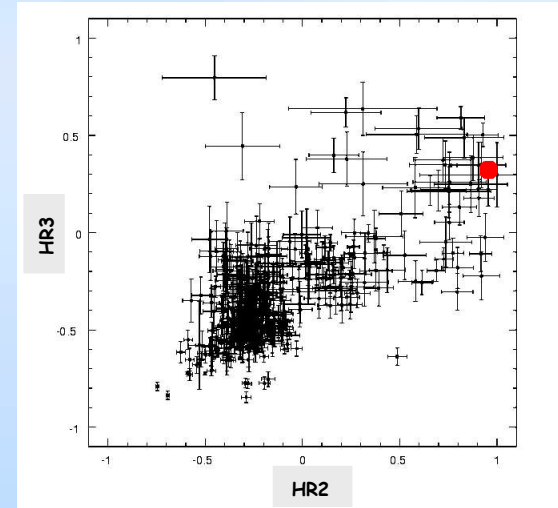
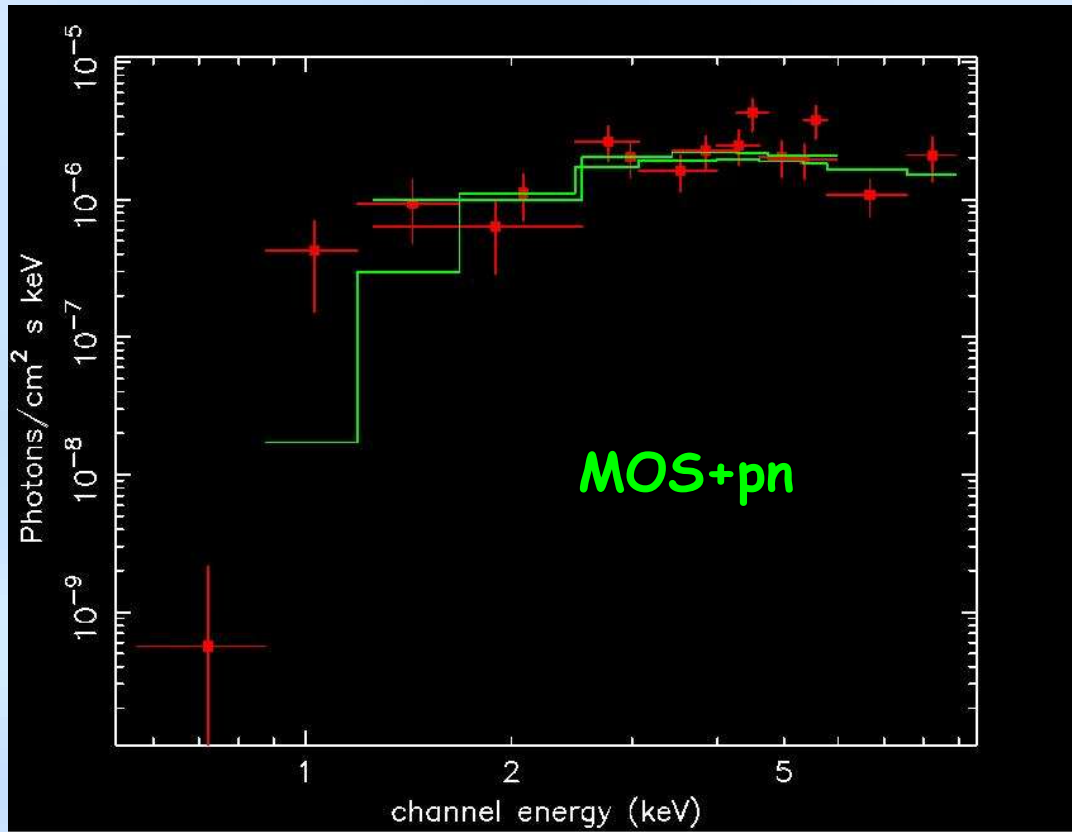
$$N_H \approx 1.5 \pm 0.5 \times 10^{22} \text{ cm}^{-2}$$

$$Fe_{\text{line}} \approx 6.34^{+0.43}_{-0.11} \text{ EW} \approx 700 \text{ eV}$$

$$L_{\text{Intr}(2-10 \text{ keV})} \approx 8 \times 10^{42} \text{ cgs}$$

Memola et al. 2006, in preparation

Newly identified absorbed AGN @OABrera



XMM X-ray spectrum

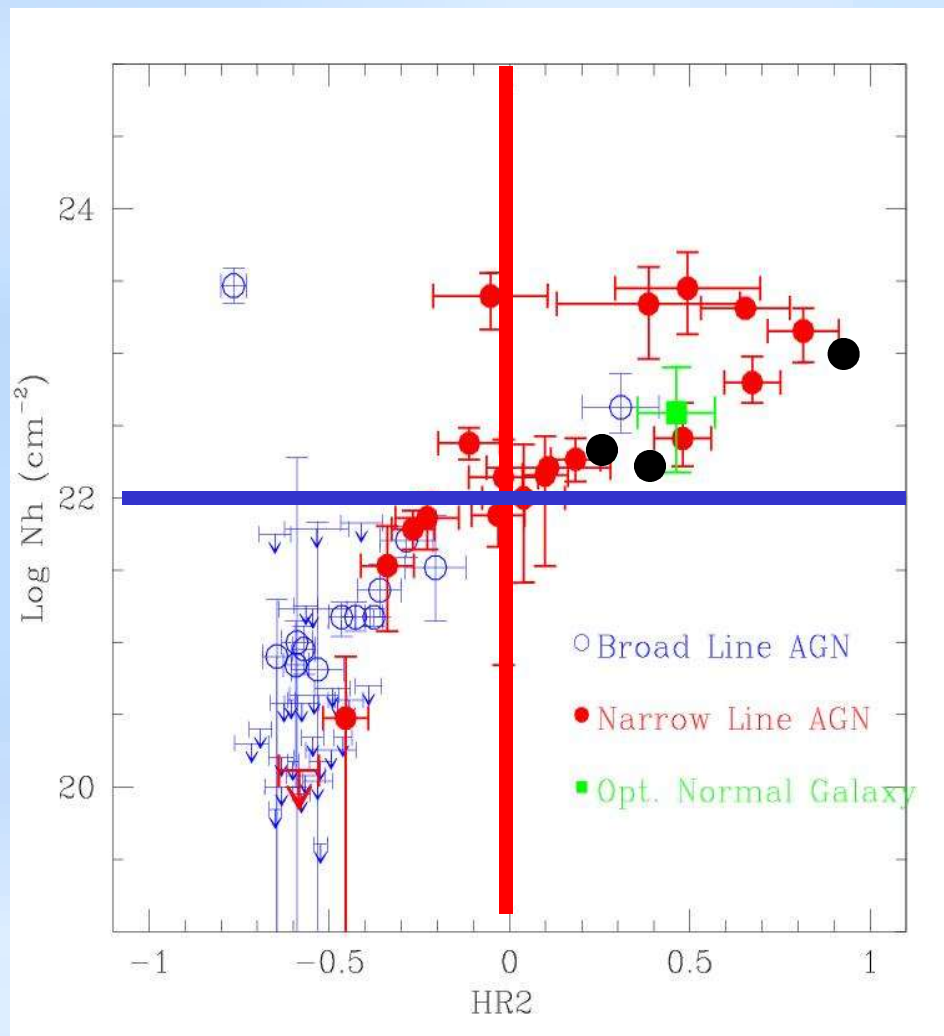
$\Gamma \approx 1.9$ (frozen)

$N_H \approx 1.1^{+0.45}_{-0.27} \times 10^{23} \text{ cm}^{-2}$

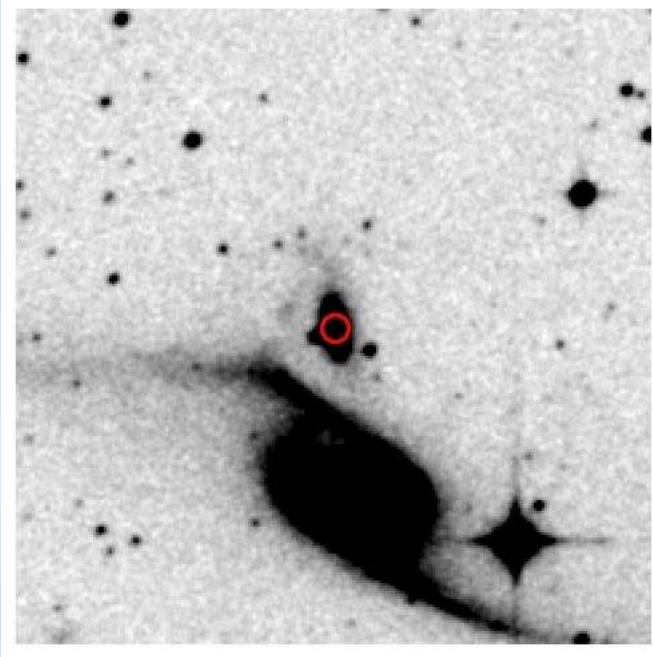
$L_{\text{Intr}(2-10 \text{ keV})} \approx 2 \times 10^{43} \text{ cgs}$

Memola et al. 2006, in preparation

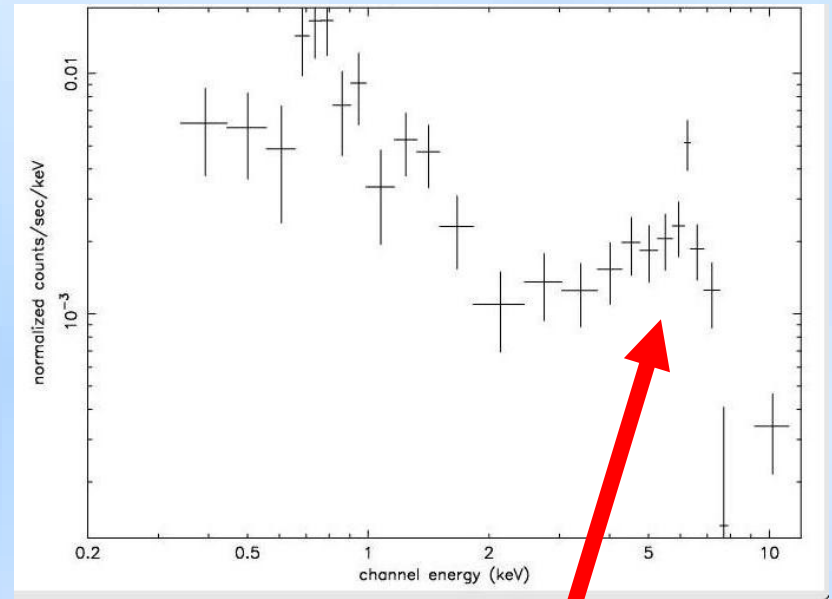
A confirmation of the criteria efficiency



An interesting example

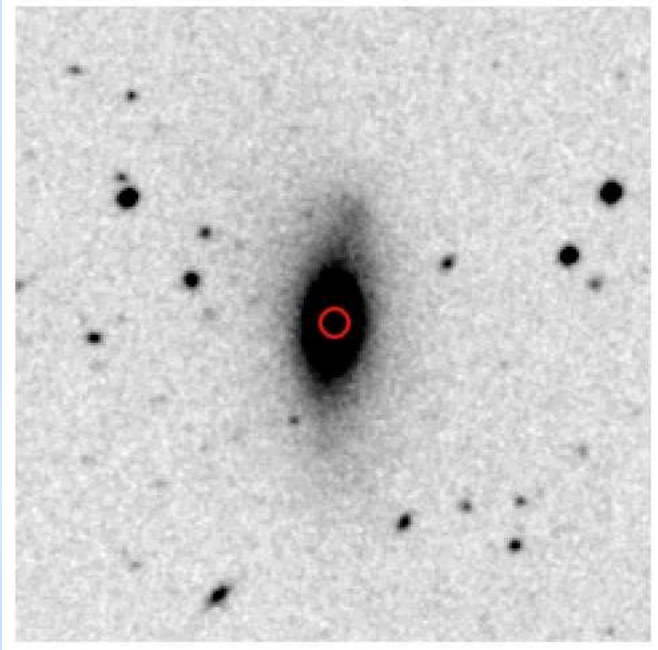


$Z=0.016$

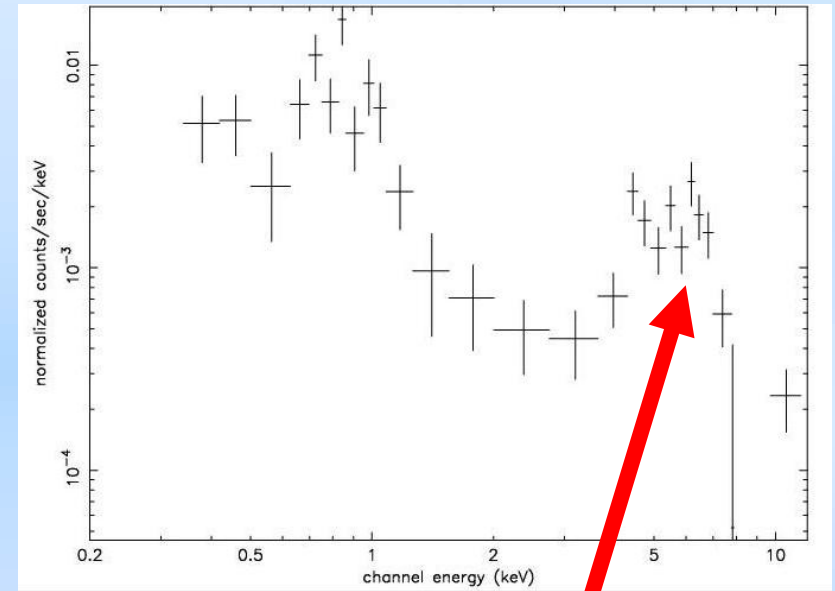


This close source shows in X-rays the typical cutoff of absorbed AGN

Another interesting example



$Z=0.021$



This close source also shows in X-rays the typical cutoff of absorbed AGN

Summary and Future



2XMM-Pilot-Sample of 54 sources



14 ID (10 literature + 4 NTT):

10 Ty2 AGN (Sy 1.8,1.9,2), 1 Sy 1.5 (?), 2 Ty2 QSOs, 1 cluster

Efficient criteria to select optically absorbed sources !!!

3 X-ray spectra analyzed: $N_H > 10^{22} \text{ cm}^{-2}$

TNG proposal (submitted) - 17 sources

Final Goals



Discover about hundred absorbed AGN
to trace the accretion history of the Universe



Where are the high Luminosity AGN with narrow lines
and how many are them? ~30% of the HBS are QSO2!

We thank ASI and INAF for partial financial support