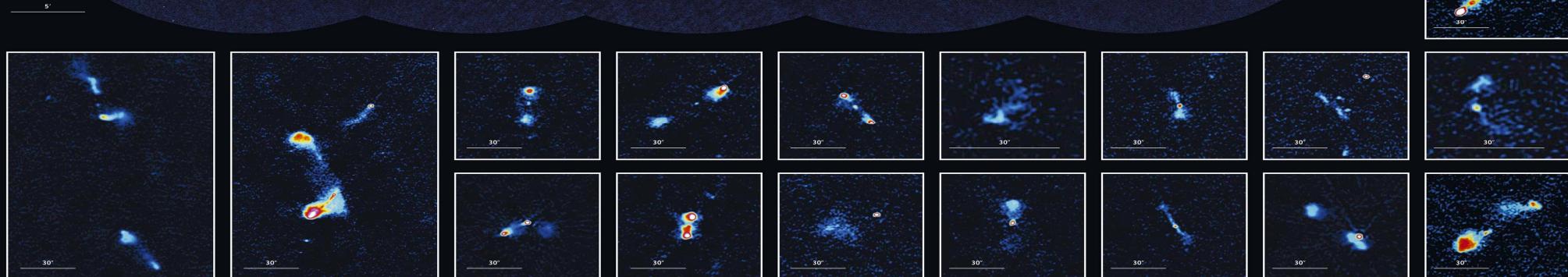


# The Radio View of the COSMOS Field

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E. Schinnerer (MPIA, Heidelberg), V. Smolcic (MPIA, Heidelberg)  
C. Carilli (NRAO, Socorro)  
& the VLA-COSMOS team

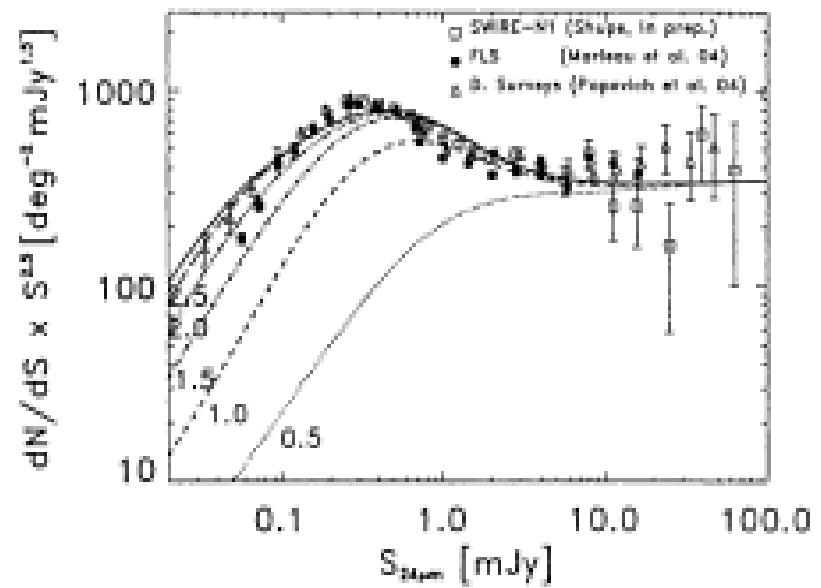


# Just to remind what COSMOS is...

- 2 sq.deg equatorial field
- HST treasury project: 640 orbits with ACS
- Worldwide collaboration providing ancillary data:  
radio (VLA), infrared (Spitzer), optical-nir (VLT,  
Subaru, NICMOS, CFHT, ...), ultraviolet (Galex),  
X-ray (XMM) + spectroscopy (Magellan,  $z$ Cosmos)

# Are radio observations important ? - I

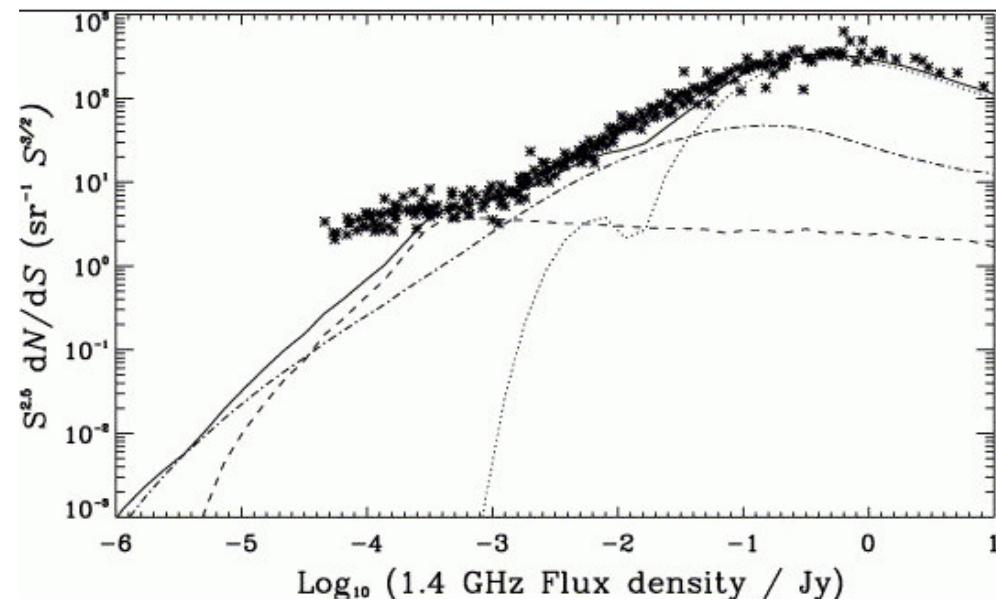
- Star Forming Galaxies
  - Radio emission unaffected by obscuration and well correlated with star-formation rate (Condon 1982)
  - Evolution of star formation rate with  $z$
  - *Redshift indicator for sub-mm galaxies without optical id*



Pozzi et al. 2005

# Are radio observations important ? -II

- AGN
  - Evolution of the faint end of the AGN radio LF
  - Large number of radio quiet AGN
  - *Deep radio observations (10.000 s) can easily detect Compton thick radio quiet AGNs*

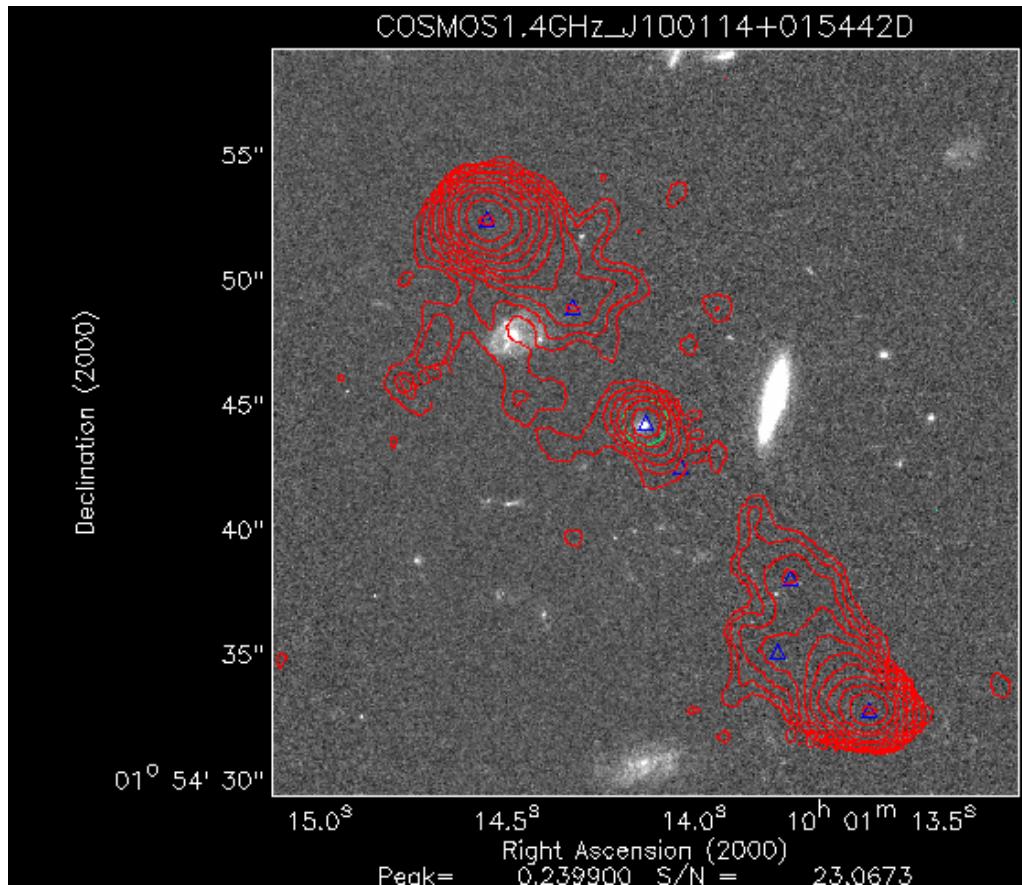


Jarvis & Rawlings 2004

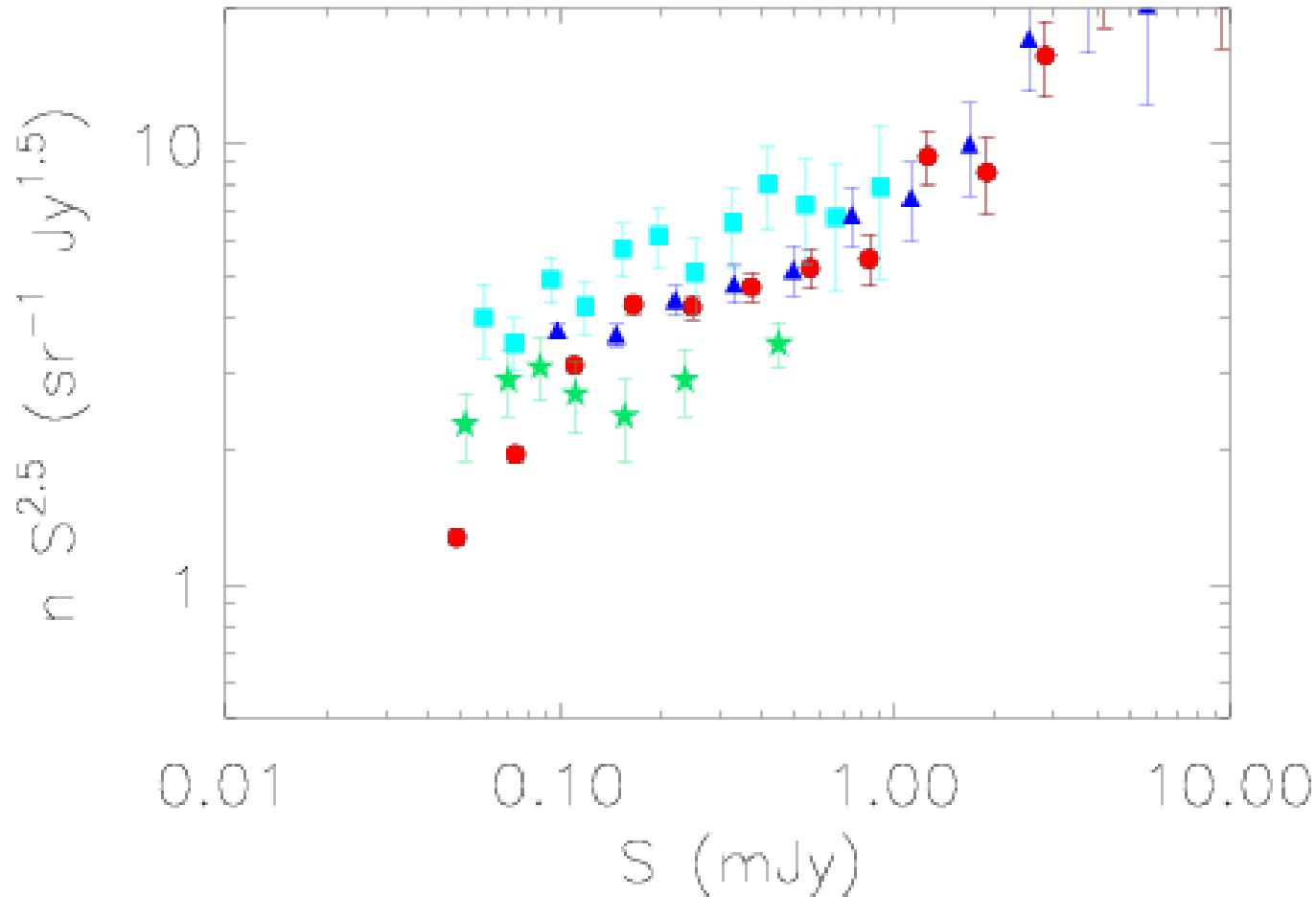
$$S_{2-10\text{keV}}(\text{Wm}^{-2}) \simeq 5 \times 10^{-16} S_{1.4}(\text{mJy})$$

# The VLA-COSMOS Survey

- Resolution 1.5 arcsec
- Sensitivity: 9-15 microJy
- Catalogue of 3643 radio sources (4.5 sigma)
- Astrometric accuracy: 100 mas



# Radio Source Counts not yet corrected for incompleteness

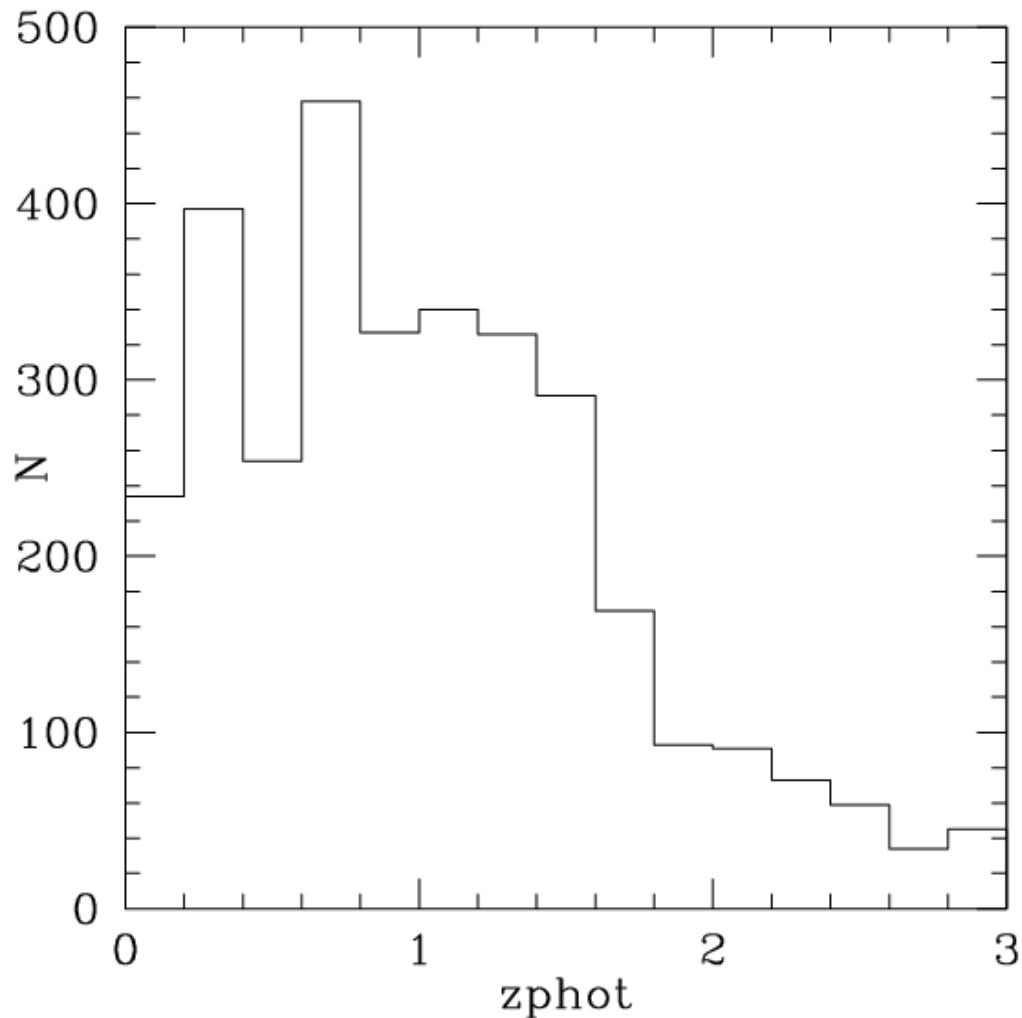


# Optical Identification

- Cross-correlation R=1.5 arcsec with optical catalog (version Jan 2006)
- 90% of radio sources optically identified

Flux interval 1.4 GHz ( $\mu$ Jy)	Early type	Late type	Starburst	Total
$S < 60$	29%	34%	37%	80%
$60 \leq S < 100$	34%	32%	34%	84%
$100 \leq S < 150$	42%	32%	26%	90%
$150 \leq S < 500$	48%	33%	19%	95%
$S \geq 500$	61%	20%	19%	95%

# Photometric Redshift



- Spurious peaks ?
- Magellan-COSMOS  
 $i(\text{AB})=23$ : 200/500  
radio sources observed
- zCOSMOS  $B(\text{AB})=25$   
compulsory targets
- Expected 1000 radio  
sources with  $z_{\text{spec}}$

# EROS

